

# Ph 406 FINAL EXAM SOLUTIONS

① 3 NUCLEON CHARGES: +, -, AND 0

$$A \equiv N_+ + N_0 + N_-$$

$$Z \equiv N_+ - N_- = \text{TOTAL CHARGE}$$

a) MASS FORMULA:  $M = A(1 - a)$

$a$  = BINDING ENERGY PER NUCLEON IN EXTENDED NUCLEAR MEDIUM

$$+ b A^{2/3}$$

LOSS OF BINDING ENERGY AT SURFACE, WHOSE AREA VARIES AS  $A^{2/3}$

$$+ d \frac{Z^2}{A^{1/3}}$$

ELECTROSTATIC ENERGY OF CHARGE  $Z$  IS SPHERE OF RADIUS  $r \sim A^{1/3}$

$$+ s \frac{(N_+ - N_-)^2 + (N_+ - N_0)^2 + (N_0 - N_-)^2}{A}$$

SYMMETRY ENERGY IN FERMI-GAS MODEL OF NUCLEUS

$$\left( + \frac{\delta}{A} \right)$$

[PAIRING ENERGY, WHICH WE IGNORE HERE]

OF THESE, ONLY THE SYMMETRY ENERGY MIGHT NOT BE 'OBVIOUS', COMPARED TO THE MASS FORMULA IN OUR UNIVERSE. (IN FACT, WE FIND A SLIGHTLY DIFFERENT FORM) BELOW..

YOU MAY WISH TO RECALL THE FERMI GAS ARGUMENT:

FOR EACH OF THE 3 NUCLEON TYPES  $N_i \sim \text{Vol } E_{f,i}^{3/2} \sim A E_{f,i}^{3/2}$

OR  $E_{f,i} \sim \left( \frac{N_i}{A} \right)^{2/3}$  = FERMI ENERGY OF TYPE  $i$

$U_i \approx N_i E_{f,i} \sim \frac{N_i^{5/3}}{A^{2/3}}$  = ENERGY (MASS) OF TYPE  $i$  IN THE NUCLEAR POTENTIAL WELL

NOW  $N_i \approx A/3$  SO WRITE  $N_i = A/3 + (N_i - A/3) = A/3 (1 + \epsilon_i)$ ,  $\epsilon_i \equiv \frac{3N_i}{A} - 1$

$$U_i \sim \frac{N_i^{5/3}}{A^{2/3}} \approx \frac{A^{5/3}}{A^{2/3}} (1 + \epsilon_i)^{5/3} \sim A \left( 1 + \frac{5}{3} \epsilon_i + \frac{1}{2} \left( \frac{5}{3} \right) \left( \frac{5}{3} \right) \epsilon_i^2 + \dots \right)$$

$$\sum_i U_i \sim A + \frac{s}{A} \sum_i (3N_i - A)^2 \quad \text{NOTICE } \sum_i \epsilon_i = 0$$

FROM THE DEFINITIONS OF  $A$  &  $Z$ ,  $N_+ = \frac{A + Z - N_0}{2}$ ,  $N_- = \frac{A - Z - N_0}{2}$

$$\text{SO } 3N_+ - A = \frac{A}{2} + \frac{3Z}{2} - \frac{3N_0}{2} = \frac{(A - 3N_0) + 3Z}{2}$$

$$3N_- - A = \frac{A}{2} - \frac{3Z}{2} - \frac{3N_0}{2} = \frac{(A - 3N_0) - 3Z}{2}$$

$$\text{SO } \sum (3N_i - A)^2 = \frac{2(A - 3N_0)^2}{4} + 2 \cdot \frac{9}{4} Z^2 + (3N_0 - A)^2 = \frac{3}{2} (A - 3N_0)^2 + \frac{9}{2} Z^2$$

SO THE SYMMETRY TERM CAN BE WRITTEN AS  $\frac{s}{A} \left( Z^2 + \frac{(A - 3N_0)^2}{3} \right)$

① b) AT FIXED A,  $M(Z, N_0) \sim M_0 + M_1 Z^2 + M_2 (A - 3N_0)^2$

THE MINIMUM IS CLEARLY AT  $Z=0$ ;  $N_0 = A/3 \Rightarrow N_+ = N_- = N_0 = A/3$

I.E.  $\beta$ -STABLE NUCLEI HAVE EQUAL NUMBERS OF ALL 3 TYPES OF NUCLEI,

AND ZERO (OR VERY SMALL) NET CHARGE

$Z \sim 0 \Rightarrow$  NO COULOMB BARRIER AGAINST FISSION, OR AGAINST

' $\alpha$ ' DECAY WHERE ' $\alpha$ ' =  $(2N_+)(2N_-)(2N_0)$

BUT ARE THESE DECAYS ALLOWED BY ENERGY CONSERVATION?

FOR  $Z=0$ ,  $N_0 = A/3$ ,  $M(A) = A(1-a) + b A^{2/3}$

FISSION:  $A \rightarrow A/2 + A/2$   $M(A/2) = \frac{A}{2}(1-a) + b \frac{A^{2/3}}{2^{2/3}}$

$M(A) - M(A/2) - M(A/2) = b A^{2/3} \left(1 - \frac{2}{2^{2/3}}\right) < 0 \Rightarrow$  FISSION IS FORBIDDEN!

IF THE ' $\alpha$ ' HAS BOUND & ENERGY GIVEN BY  $M(\alpha) = 6(1-a) + b 6^{2/3}$

AND  $M(A-6) = (A-6)(1-a) + b(A-6)^{2/3}$

$\approx A^{2/3} \left(1 - \frac{6}{A}\right)^{2/3} \approx A^{2/3} \left(1 - \frac{2}{3} \frac{6}{A}\right)$

$\approx (A-6)(1-a) + b A^{2/3} - \frac{4b}{A^{1/3}}$

$M(A) - M(\alpha) - M(A-6) = b \left(\frac{4}{A^{1/3}} - 6^{2/3}\right) < 0$  FOR  $A > 1$

$\Rightarrow$  ' $\alpha$ ' DECAY FORBIDDEN ALSO

$\Rightarrow$  HIGH A NUCLEI ARE STABLE IN THIS UNIVERSE!

IN FACT IF TWO NUCLEI COLLIDE, IT IS ENERGETICALLY FAVORABLE THAT THEY COALESCE

$\Rightarrow$  'STARS' CONSIST OF A SINGLE GIANT NUCLEUS, WITH  $Z \sim 0$

$\Rightarrow$  THEY DON'T GLOW DUE TO ANY NUCLEAR PROCESS (A BIT LIKE A NEUTRON STAR)

SO LIFE IS BAD IN THIS UNIVERSE FOR VARIOUS REASONS:

NO STARS

NO CHEMISTRY AS  $Z \approx 2$  ALWAYS

$\Rightarrow$  NO LIFE...

(2) a) For a rotational excitation,  $E = \frac{1}{2} I \omega^2$  where  $\vec{J} = I \vec{\omega}$

so  $E = \frac{|\vec{J}|^2}{2I} = \frac{J(J+1)}{2I}$  NOTING EIGENVALUES OF  $\vec{J}^2$

NOW A ROTATING SPHERE (OR EVEN ELLIPSOID) IS SYMMETRIC UNDER THE PARITY TRANSFORMATION  $\vec{r} \rightarrow -\vec{r} \Rightarrow P = +$  ONLY. BUT  $P = (-1)^J$  FOR ROTATIONS, SO HAVE EVEN J ONLY

HENCE  $J^P = 0^+, 2^+, \dots (2i)^+$  IS EXPECTED SEQUENCE OF STATES

ALSO, EXPECT	J	0	2	4	6	8
	$2I \cdot E = J(J+1)$	0	6	20	42	72

THIS SEQUENCE OF ENERGIES IS FAIRLY WELL MATCHED TO THE DATA, WITH  $2I \cdot 100 \sim 6 \text{ (KEV)}^{-1}$  OR  $I \sim 0.03 \text{ (KEV)}^{-1}$

COMPARE  $I \sim \frac{2}{5} M R^2$  FOR  $A = 170$

$M \sim 170 \cdot 930 \text{ MeV}$

$R \sim 1.1 A^{1/3} \text{ FERMI} \sim 6 \text{ F}$

$I \sim (0.4) (170) (930) 36 \text{ MeV} \cdot \text{FERMI}^2$

NOW  $1 \text{ FERMI} \sim \frac{1}{200} \text{ MeV}$

SO  $I \sim (0.4) \frac{(170)(930)(36)}{(200)(200) \text{ MeV}} \sim \frac{50.0}{\text{MeV}} \sim \frac{50.0}{1000 \text{ KEV}} \sim \underline{0.05 \text{ (KEV)}^{-1}}$

IN FAIR AGREEMENT WITH THE DATA (THO  $I_{\text{DATA}} < I_{\text{MODEL}}$  SUGGESTS THE WHOLE NUCLEUS IS NOT ROTATING AS A RIGID BODY....)

EQUIVALENTLY  $I_{\text{MODEL}} \sim \frac{50}{1 \text{ MeV}} \cdot \frac{200 \text{ MeV} \cdot \text{FM}}{1} \sim 10^4 \text{ FERMI}^2$

$I_{\text{DATA}} \sim 5 \times 10^3 \text{ FERMI}^2$

② b)  $^{15}_7\text{N}$  LEVELS:

IN THE SHELL MODEL THE GROUND STATE HAS NEUTRONS FILLING THE  $1S_{1/2}$ ,  $1P_{3/2}$  &  $1P_{1/2}$  SHELLS,

PROTONS FILLING THE  $1S_{1/2}$  &  $1P_{3/2}$  SHELLS, BUT ONE VACANCY IN THE  $1P_{1/2}$ .

THE UNPAIRED PROTON DETERMINES THE GROUND STATE TO BE  $1/2^-$

1ST EXCITED STATE IS  $5/2^+$ , CONSISTENT WITH THE UNPAIRED PROTON EXCITED TO THE  $1d_{5/2}$  SHELL

2ND EXCITED STATE IS  $1/2^+$ : NOW THE UNPAIRED PROTON IS IN THE NEXT HIGHER SHELL,  $2S_{1/2}$

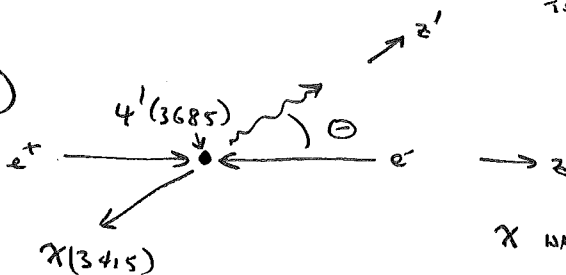
3RD EXCITED STATE IS  $3/2^-$ ; IF THE UNPAIRED PROTON WERE IN THE NEXT HIGHER SHELL,  $1d_{3/2}$ , WOULD EXPECT  $3/2^+$ !

INSTEAD, COULD BE THAT A PROTON FROM THE  $1P_{3/2}$  SHELL IS EXCITED INTO THE  $1P_{1/2}$  FILLING THE LATTER. THEN THE UNPAIRED  $1P_{3/2}$  PROTON DETERMINES THE STATE AS  $3/2^-$ .

4TH EXCITED STATE IS  $5/2^+$  COULD BE THAT A NEUTRON FROM THE FILLED  $1P_{1/2}$  SHELL IS EXCITED TO THE  $1d_{5/2}$  THEN THE UNPAIRED PROTON & NEUTRON IN THE  $1P_{1/2}$  TOGETHER GIVE  $3P=0^+$ , AND THE EXCITED NEUTRON DETERMINES THE STATE TO BE  $5/2^+$

5TH EXCITED STATE IS  $3/2^+$ . THE UNPAIRED PROTON IN THE  $1P_{1/2}$  IS EXCITED TO THE  $1d_{3/2}$ , ...

③ a)



$e^+e^-$  ANNIHILATION  $\Rightarrow J_2 = \pm 1$  ONLY FOR THE  $\gamma(3685)$

X HAS  $J=0$  BY HYPOTHESIS

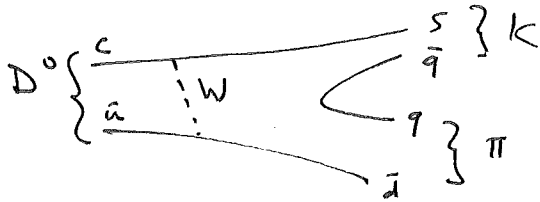
$\gamma$  HAS  $J_2' = \pm 1$  ONLY, AS IS REAL PHOTON

$\Rightarrow$  TOTAL  $J_2' = \pm 1$  ONLY

SO ANGULAR DISTRIBUTION IS GIVEN BY PROJECTING  $J_2 = \pm 1$  ONTO  $J_2' = \pm 1$

$$\begin{aligned}
 I_{\theta} \propto & |D_{11}^1|^2 + |D_{1-1}^1|^2 + |D_{-11}^1|^2 + |D_{-1-1}^1|^2 \\
 \sim & \left| \frac{1 + \cos \theta}{2} \right|^2 + \left| \frac{1 - \cos \theta}{2} \right|^2 \sim \underline{\underline{1 + \cos^2 \theta}}
 \end{aligned}$$

(3) b)



USE ISOSPIN ANALYSIS. BUT WITH SOME CAUTION, ISOSPIN IS NOT CONSERVED IN THE WEAK INTERACTION  $c\bar{u} \rightarrow s\bar{d}$  !

INDEED  $c\bar{u}$  HAS  $I_3 = -1/2$  (SINCE A  $\bar{u}$  HAS  $I_3 = -1/2$ )

WHILE  $s\bar{d}$  HAS  $I_3 = +1/2$

BUT ISOSPIN IS CONSERVED IN THE STRONG INTERACTION  $s\bar{d} \rightarrow s\bar{q}q\bar{d}$

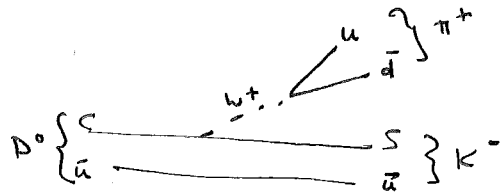
SO WE CAN WRITE  $|s\bar{d}\rangle \rightarrow c_1 |s\bar{u}\rangle |u\bar{d}\rangle + c_2 |s\bar{d}\rangle |d\bar{d}\rangle$   
 $K^- \quad \pi^+ \quad \quad \quad \bar{K}^0 (-\frac{1}{\sqrt{2}}\pi^0)$  NOTING  $(\pi^0) \propto (u\bar{u}) - (d\bar{d}) / \sqrt{2}$

OR  $|\frac{1}{2}, \frac{1}{2}\rangle \rightarrow \sqrt{\frac{2}{3}} |\frac{1}{2}, -\frac{1}{2}\rangle |1, 1\rangle + (-\frac{1}{\sqrt{3}}) (\sqrt{\frac{1}{2}}) |\frac{1}{2}, \frac{1}{2}\rangle |1, 0\rangle$

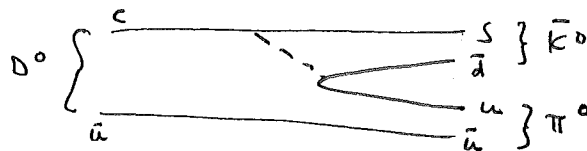
USING THE CLEBSCH-GORDAN CRIB SHEET FOR  $c_1$  &  $c_2$ .

THUS  $\frac{\Gamma(D^0 \rightarrow K^- \pi^+)}{\Gamma(D^0 \rightarrow \bar{K}^0 \pi^+)} = \frac{|\sqrt{\frac{2}{3}}|^2}{|\frac{1}{\sqrt{6}}|^2} = 4$

ANOTHER CABIBBO-FAVORED DIAGRAM IS

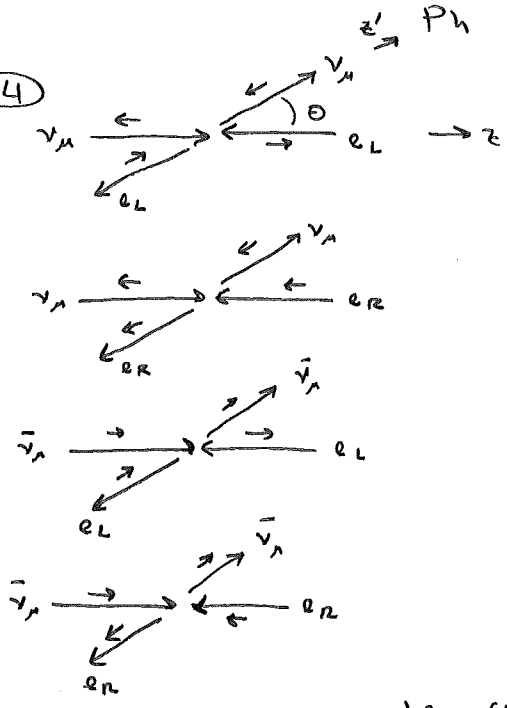


AND ALSO



THESE MAY WELL HAVE COMPARABLE STRENGTHS, AND SO TEND TO REDUCE THE RATIO OF BRANCHING RATIOS TO LESS THAN 4....

(4)



CASES 1 & 4 HAVE  $J_z \in J_z'$  BOTH 0, SO EXPECT ISOTROPIC ANGULAR DISTRIBUTION

CASE 2 WAS  $J_z = -1, J_z' = -1$

CASE 3 WAS  $J_z = 1, J_z' = 1$

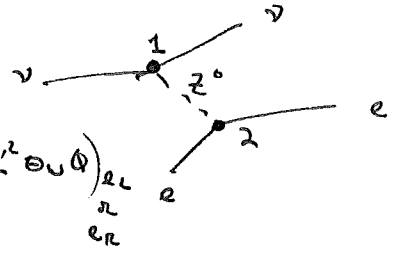
SO EXPECT ANGULAR DIST  $\sim |D_{11}|^2 \sim \left| \frac{1 + \cos\theta}{2} \right|^2$

INTEGRATION OVER ANGLE,  $\int_{-1}^1 d\cos\theta \left| \frac{1 + \cos\theta}{2} \right|^2 = \frac{1}{4} (2 + \frac{2}{3}) = \frac{2}{3}$

COMPARED TO  $\int_{-1}^1 d\cos\theta = 2$ .

I.E. CASES 2 & 3 HAVE  $\frac{1}{3}$  THE CROSS SECTION OF 1 & 4

TO GET THE VERTEX FACTORS, CONSIDER THE DIAGRAM



AMPLI  $\sim \frac{g}{\sqrt{2}} (\bar{I}_3 - \sin^2\theta_W Q)_{\nu \nu} \left( \frac{1}{M_Z^2} \right) \frac{g}{\sqrt{2}} (\bar{I}_3 - \sin^2\theta_W Q)_{e_L, e_R}$   
 ↑  
 PROPAGATOR

$\sim \frac{G}{\sqrt{2}} \left( \frac{1}{2} \right) \cdot \begin{cases} -\frac{1}{2} - (\sin^2\theta_W)(-1) & \text{FOR } e_L \\ 0 - \sin^2\theta_W(-1) & \text{FOR } e_R \end{cases}$

USING  $\bar{I}_3(\nu \text{ or } \bar{\nu}) = \frac{1}{2}$   
 $\bar{I}_3(e_L) = -\frac{1}{2}$   
 $\bar{I}_3(e_R) = 0$   
 ↑  
 WEAK ISOSPIN

$\sigma \sim \sum_{e_L, e_R} (\text{AMPLI})^2 \cdot (\text{ANGULAR FACTOR}) \cdot (\text{CM ENERGY})^2$   
 ↑ BY DIMENSIONAL ANALYSIS, SINCE  $\sigma \sim M^2$   
 $G \sim \frac{1}{M^2}$ , AND  $(\text{AMPLI})^2 \sim G^2$

SO  $\sigma \sim G E_{cm}^2 \left\{ \begin{array}{l} \left( \frac{1}{2} - \sin^2\theta_W \right)^2 + \frac{1}{3} \sin^2\theta_W \quad \text{FOR } \nu_A \text{ or } \bar{\nu}_A \\ \frac{1}{3} \left( \frac{1}{2} - \sin^2\theta_W \right)^2 + \sin^2\theta_W \quad \text{FOR } e_L \text{ or } e_R \end{array} \right.$

[ DETAILED CALCULATION SHOWS THERE IS AN OVERALL FACTOR OF  $\frac{4}{\pi}$  MISSING FROM THE QUICK ESTIMATE. ]